Solar and Terrestrial Radiation

Energy in the Atmosphere

Ch. 2: Energy in the Atmosphere

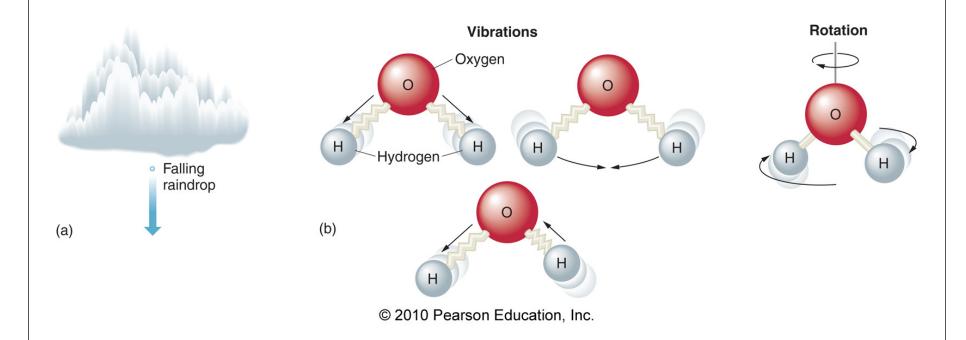
- Heat, Energy, Temperature
 - → Definitions
 - → Temperature Scales
- Heat Transfer
 - → Conduction, Convection, Radiation
- Electromagnetic Radiation (EMR)
 - → Types: Ultraviolet, Visible, Infrared
 - → Blackbody Radiation

Lecture Topics, cont.

- Seasons
 - **→ Earth's Axial Tilt**
 - → Solstices and Equinoxes
 - → Solar Angle and Daylight Length

Heat and Temperature

- Kinetic Energy = energy of motion
 - → Higher speed → Higher KE



Heat



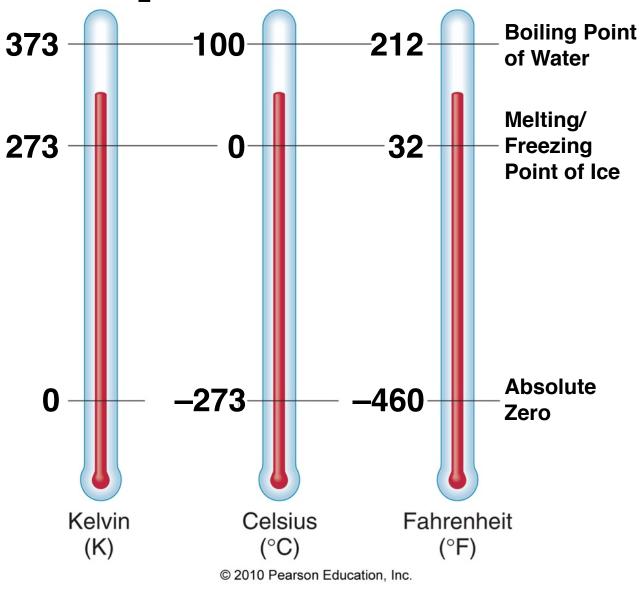
Cumulative total kinetic energy of all molecules in a substance

Temperature



Proportional to average kinetic energy of individual molecules in a substance

Temperature Scales



Temperature Conversion

$$K = C + 273$$

$$C = \frac{5}{9} (F - 32)$$

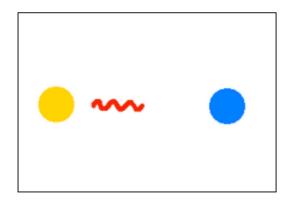
$$F = \frac{9}{5}C + 32$$

Heat Transfer

- Heat flows from regions of high heat content to regions of low heat content
 - → Usual direction: high temperature regions to low temperature regions
- Conduction: heat spread by collisions between molecules
 - → Slow!
 - → Air is a poor conductor/is a good insulator

Randomly moving air molecules

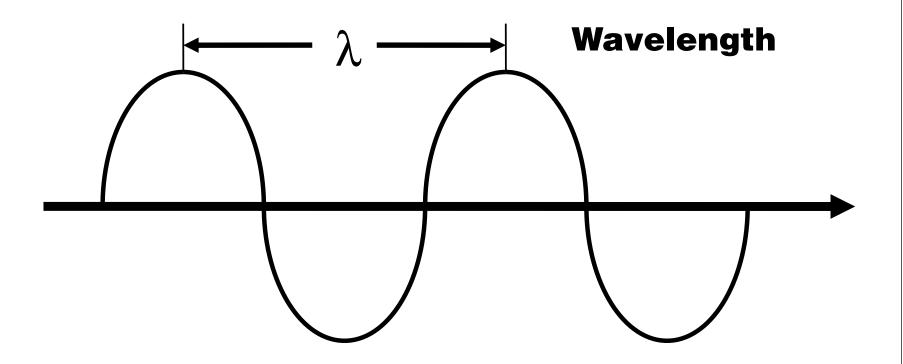
- Convection: heat transport via bulk movement of air
 - → Faster than conduction; as fast as the wind
 - → "Convection": vertical movement
 - → "Advection": horizontal movement



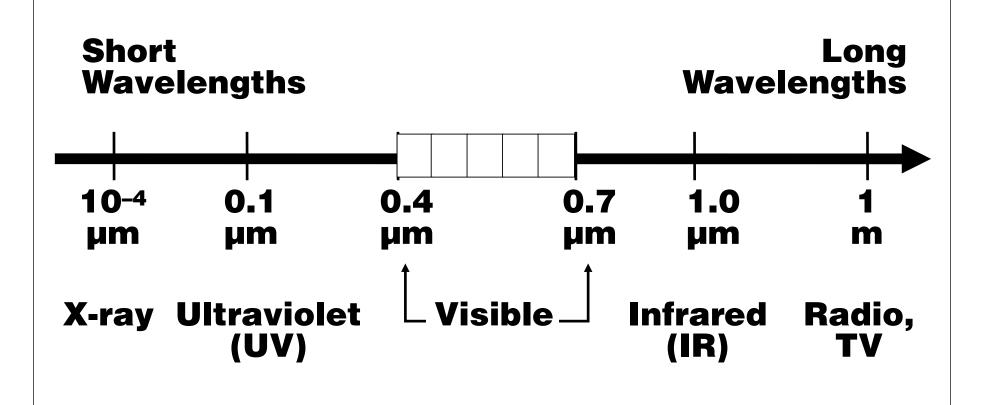
- Radiation: energy propagating through space in the form of electromagnetic radiation
 - → Moves as fast as speed of light
 - → Can travel through a vacuum

Electromagnetic Radiation

Radiation energy propagates as a wave, in energy packets called photons

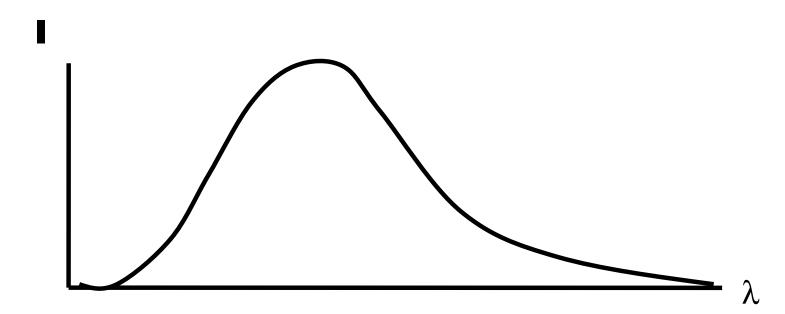


Types of EMR



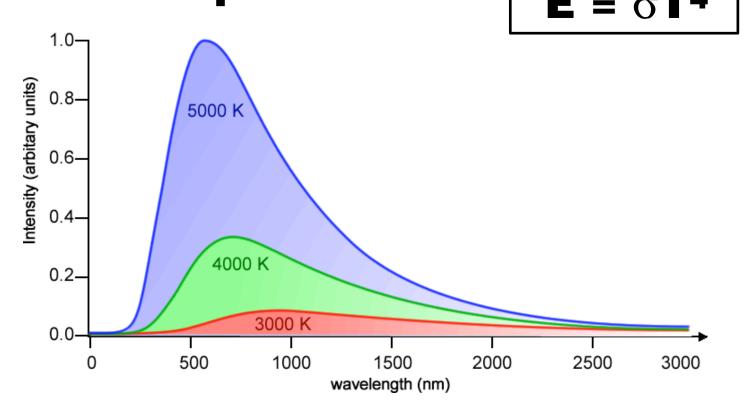
Blackbody Radiation

- Blackbody: an object that absorbs all wavelengths of EMR
 - → And emits all wavelengths of EMR (Kirchoff's Law)



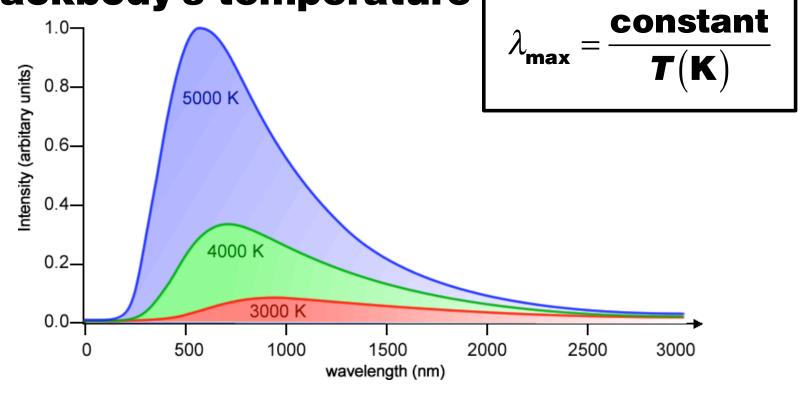
Stefan-Boltzmann Law

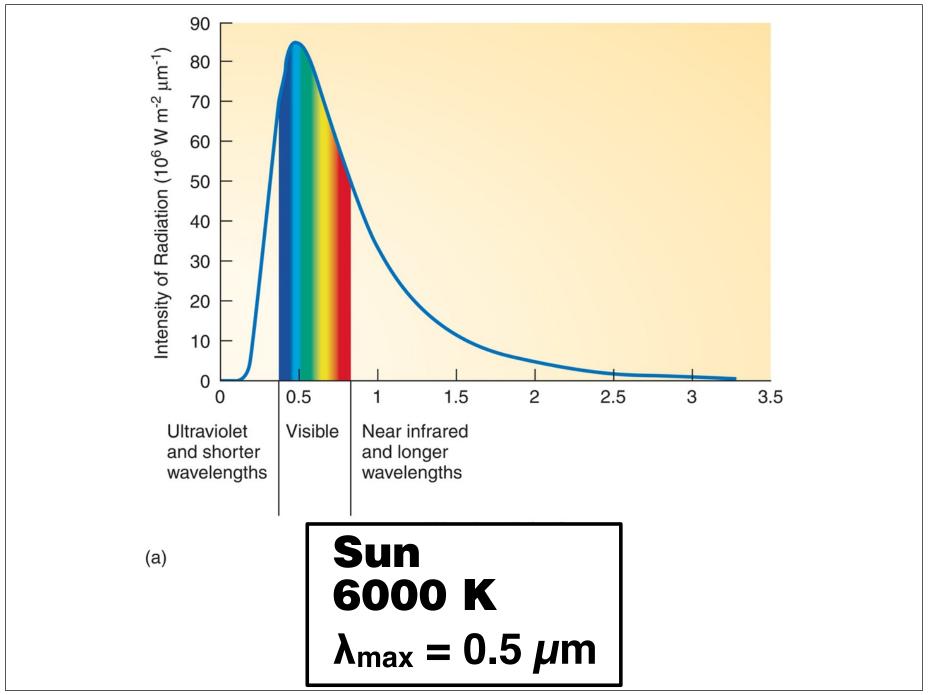
 Radiation power flux emitted by a blackbody is directly proportional to its temperature

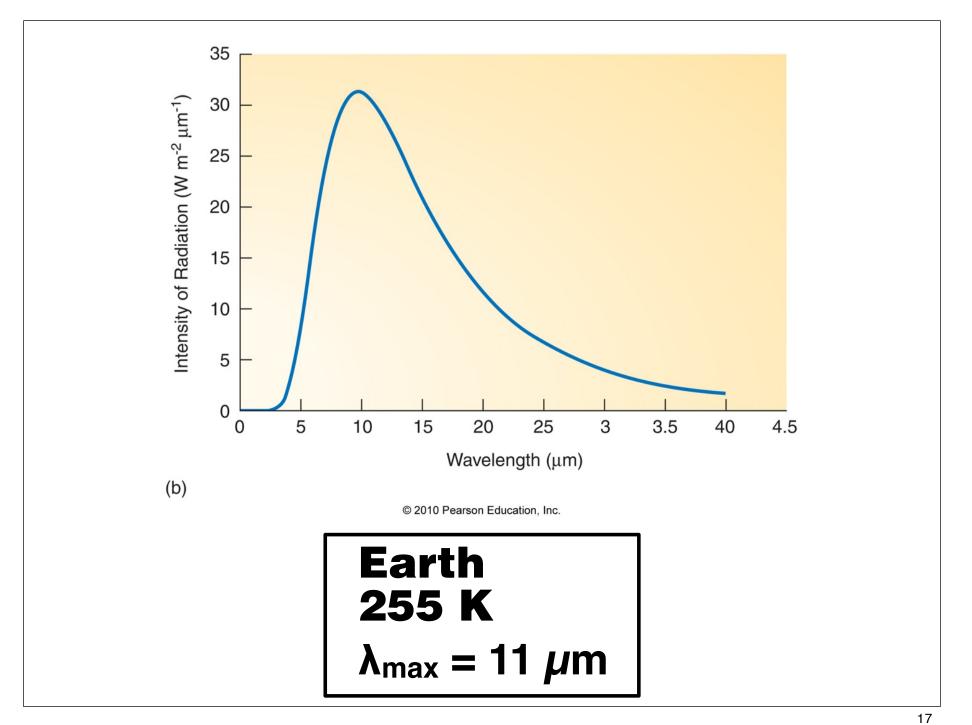


Wien's Law

• A blackbody emits EMR such that the wavelength of maximum intensity (λ_{max}) is inversely proportional to the blackbody's temperature

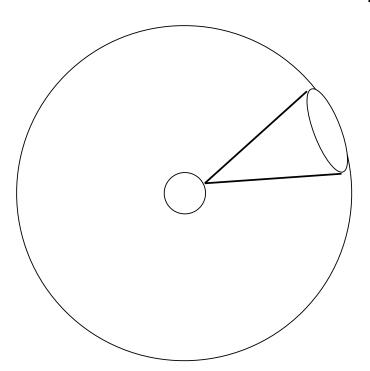


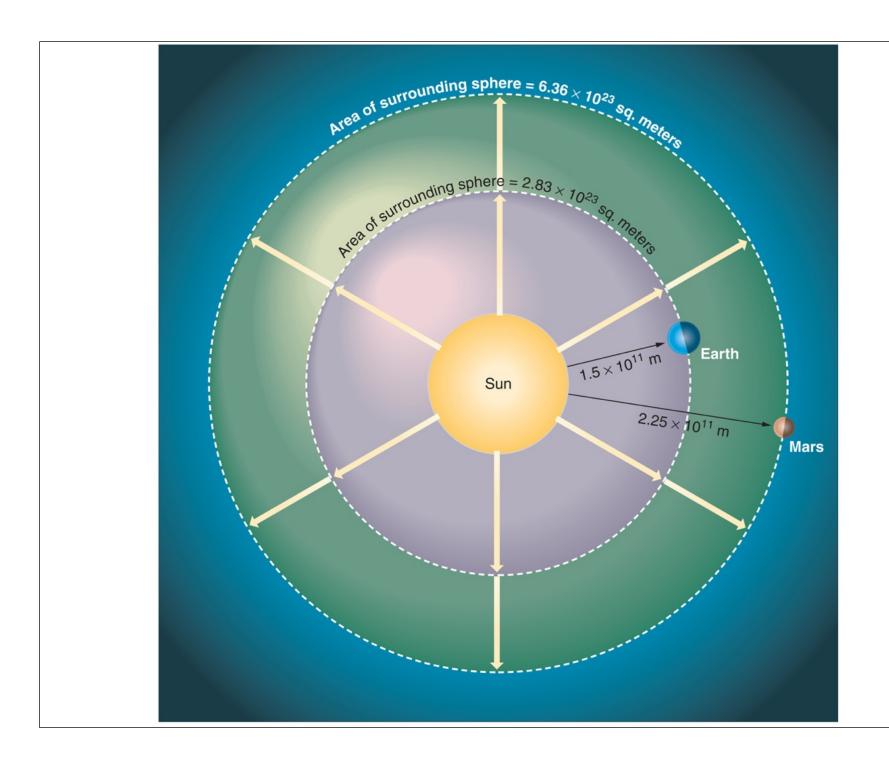




Inverse Square Law

 Radiation power flux from an object decreases as the square of the distance from the object





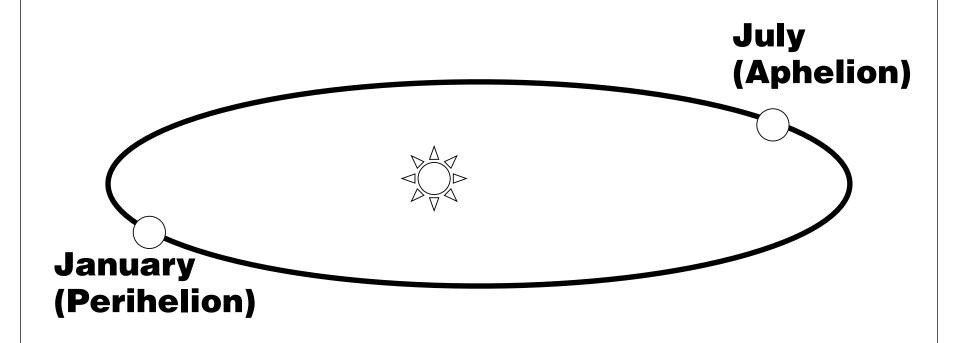
Insolation and Earth Seasons

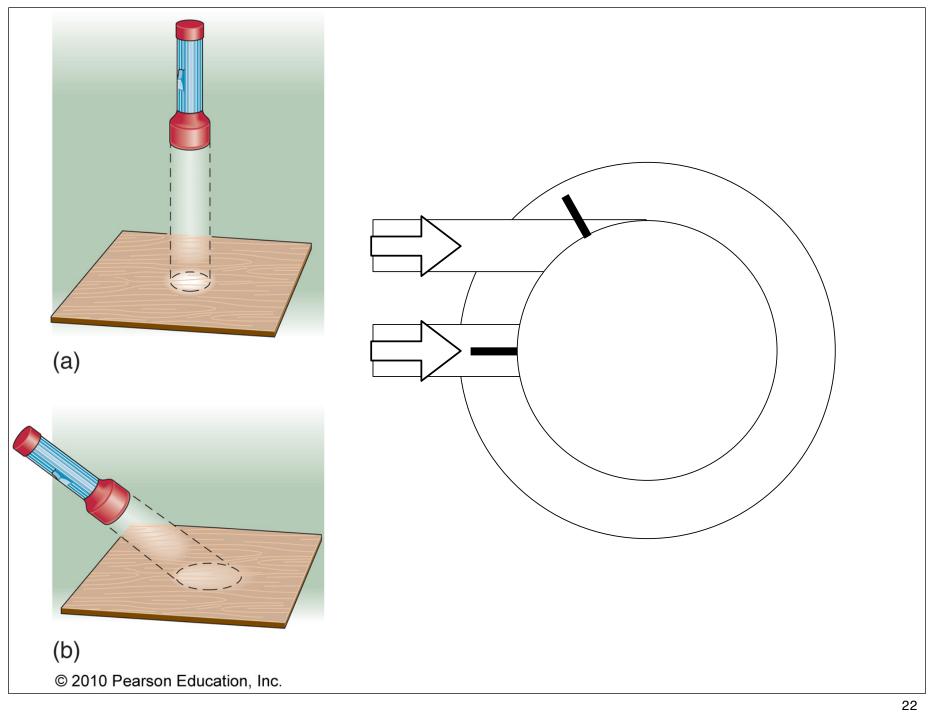
Insolation: cumulative amount of solar radiation energy received by the Earth's surface in one day

Intensity Duration

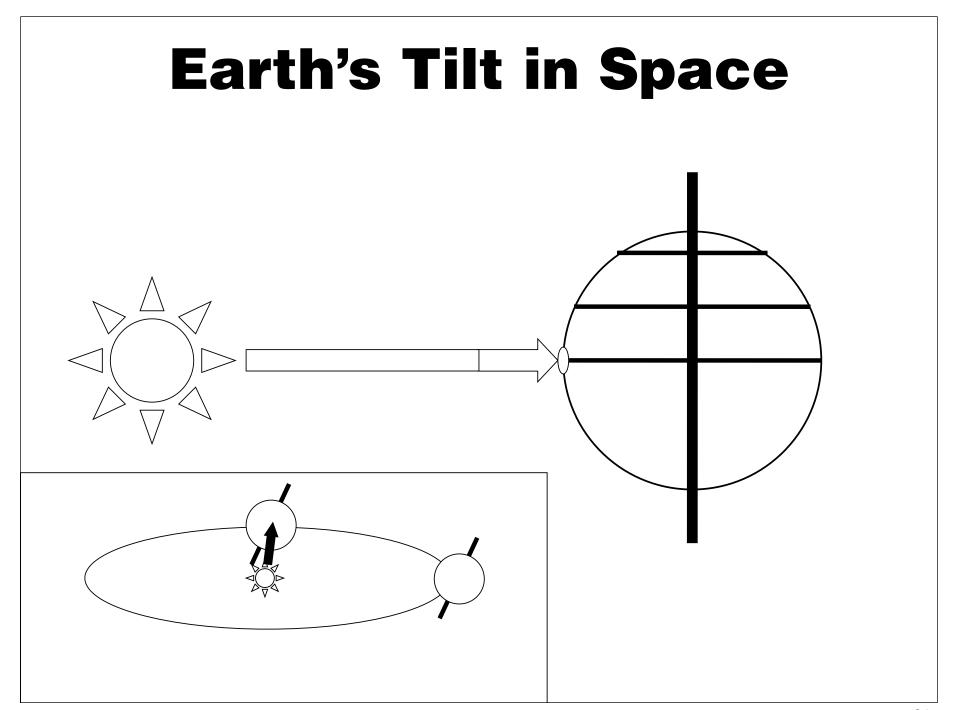
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- Seasons are regular, annual variations in temperature and/or climate
 - → Results from variation of insolation during the course of a year

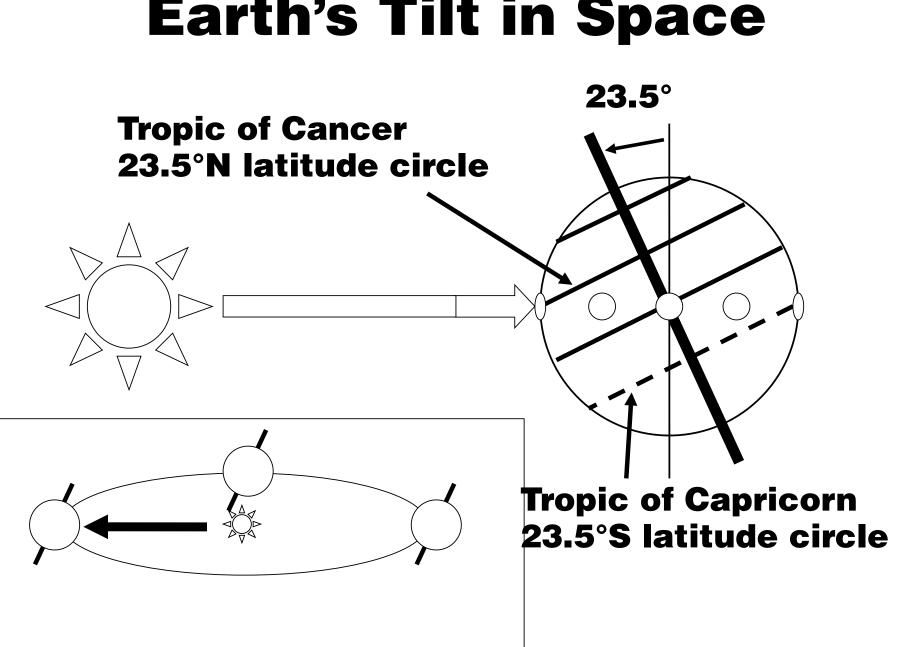




Earth's Tilt in Space 23.5°

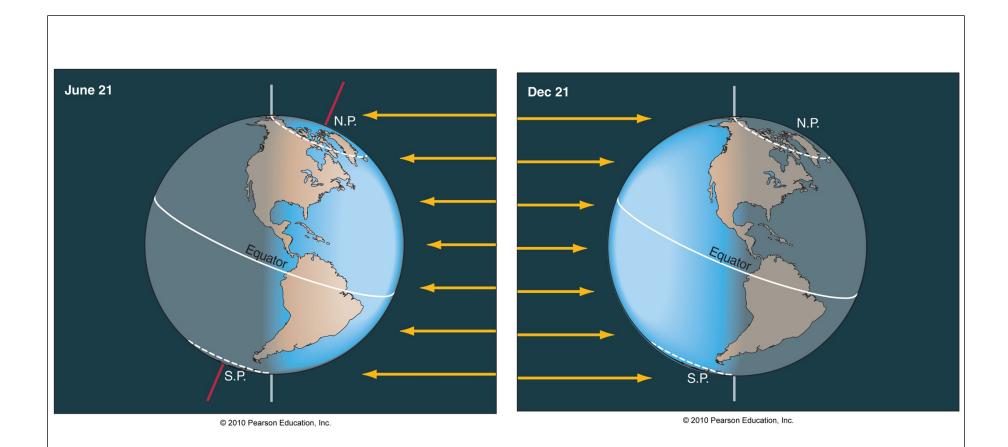


Earth's Tilt in Space



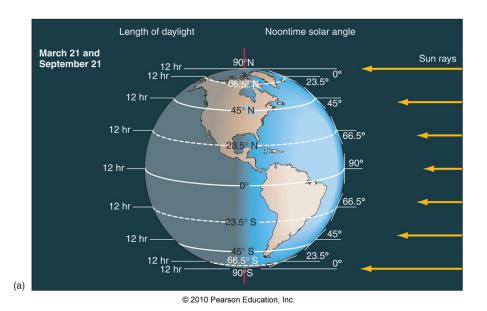
Solstices

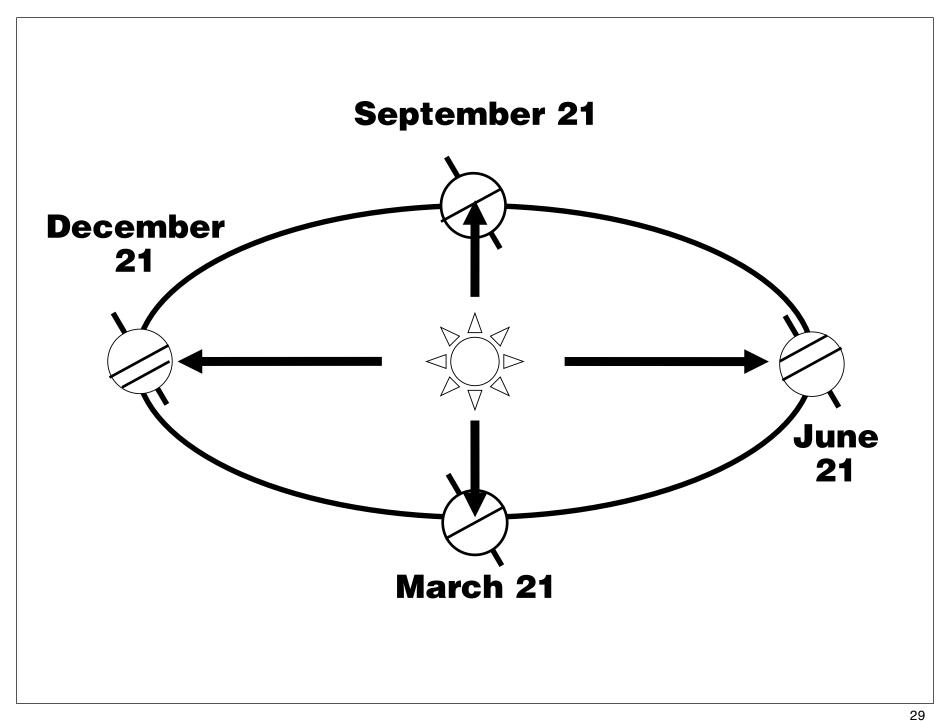
- North/South Poles are pointed most toward or away from the Sun
- Subsolar point is Tropic of Cancer or Capricorn
 - → June Solstice: ~June 21, subsolar point Tropic of Cancer
 - → December Solstice: ~December 21, subsolar point Tropic of Capricorn



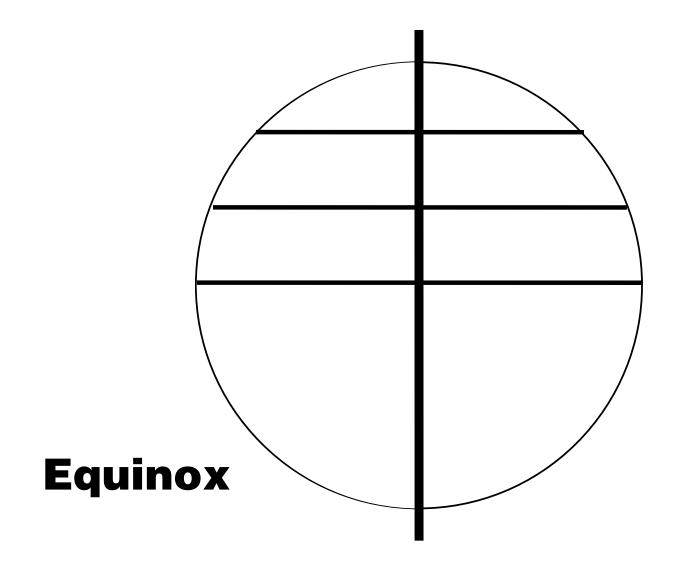
Equinoxes

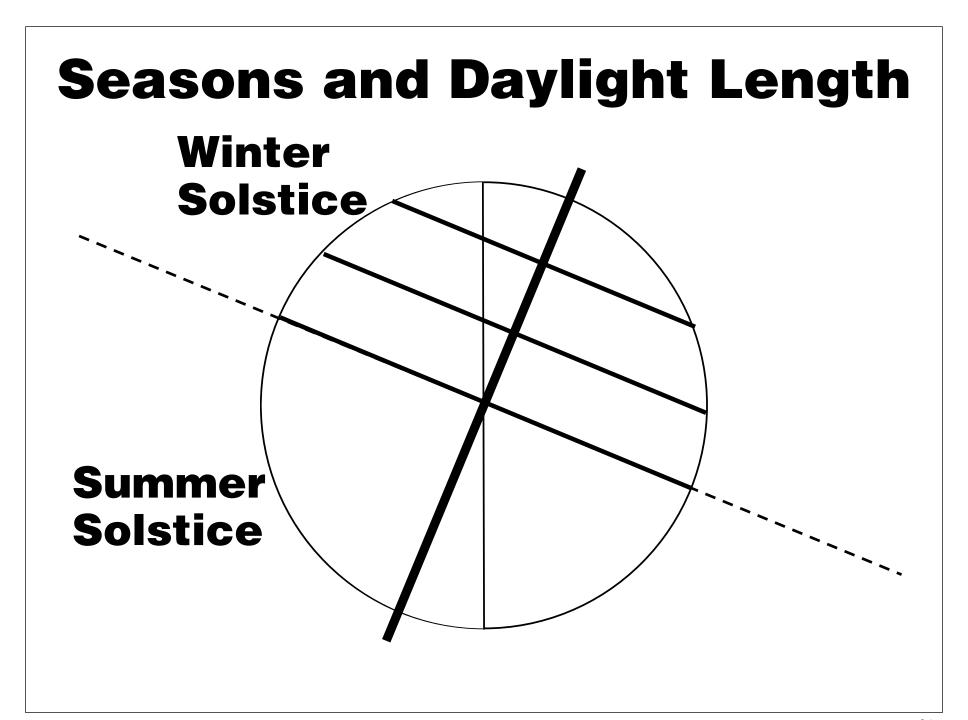
- Subsolar point = 0° latitude (the Equator)
- Directly between solstices
- 12 hours daylight, 12 hours night





Seasons and Daylight Length





Seasons and Daylight Length

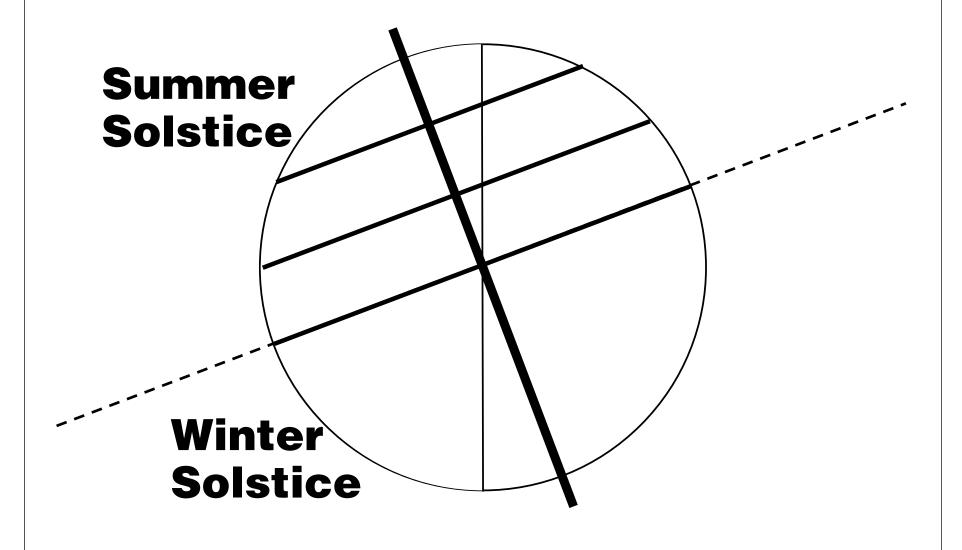


TABLE 2-2 Variations in Solar Angle and Daylength

	Solar Angle at Noon	Length of Day	Total Radiation for Day (Megajoules/m²)
December 21			
Winnipeg (50 °N)	16.5°	7 hr, 50 min	7.1
Austin (30 °N)	36.5°	10 hr, 04 min	18.6
June 21			
Winnipeg (50 °N)	63.5°	16 hr, 10 min	44.5
Austin (30 °N)	83.5°	13 hr, 56 min	43.9

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